

**NUCLEAR DATA AND MEASUREMENTS SERIES**

**ANL/NDM-113**

**Resonance Effects in Neutron Scattering Lengths**

by

J.E. Lynn

June 1989

**ARGONNE NATIONAL LABORATORY,  
ARGONNE, ILLINOIS 60439, U.S.A.**

# NUCLEAR DATA AND MEASUREMENTS SERIES

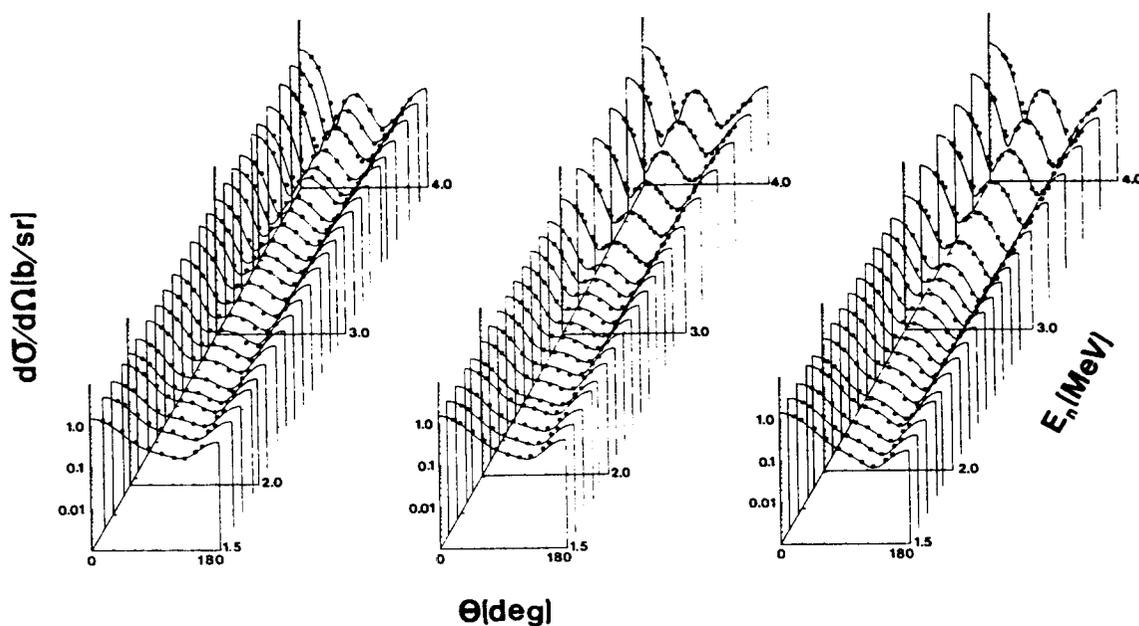
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ARGONNE NATIONAL LABORATORY, ARGONNE, ILLINOIS

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RESONANCE EFFECTS  
IN NEUTRON SCATTERING LENGTHS\*

by

J. E. Lynn†

June 1989

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NEUTRON SCATTERING LENGTHS. Neutron scattering lengths and nuclear effects giving rise to their variation are discussed.

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RESONANCE EFFECTS  
IN NEUTRON SCATTERING LENGTHS\*

by  
J. E. Lynn†

Abstract

The nature of neutron scattering lengths is described and the nuclear effects giving rise to their variation is discussed. Some examples of the shortcomings of the available nuclear data base, particularly for heavy nuclei, are given. Methods are presented for improving this data base, in particular for obtaining the energy variation of the complex coherent scattering length from long to sub-Ångstrom wave lengths from the available sources of slow neutron cross section data. Examples of this information are given for several of the rare earth nuclides. Some examples of the effect of resonances in neutron reflection and diffraction are discussed. This report documents a seminar given at Argonne National Laboratory in March 1989.

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# RESONANCE EFFECTS IN NEUTRON SCATTERING LENGTHS

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## 1. Introduction

To a nuclear physicist surveying the opportunities now being opened up in the field of neutron scattering techniques for condensed matter research, particularly by the intense spallation neutron sources that can operate well into the sub-Ångström region, it is clear that the readily accessible nuclear data base for this work is far from satisfactory. It is my purpose in this seminar to describe the current situation in the data base as I see it, and in particular to discuss the neutron coherent scattering lengths of the heavy nuclides, which often have strong energy dependences and imaginary components owing to nuclear resonances in their cross-sections.

The solid-state physicist primarily wishes to regard the nuclei off which his neutrons scatter as tiny impenetrable spheres, the Fermi pseudo-potentials, which are only characterized by their effective "radius", the scattering length. The fact that they are much more than that is already apparent from surveying the neutron scattering properties of the very lightest nuclides. First of all, we have the strong isotope variation; examples are

$${}^1\text{H} \quad a_{\text{coh}} = -1.87 \text{ fm} \quad (-0.187 \cdot 10^{-12} \text{ cm})$$

$${}^2\text{H} \quad a_{\text{coh}} = 4.45 \text{ fm}$$

( $a$  denotes free atom scattering length). The fact that the proton scattering length is negative also shows the inadequacy of the particle picture. Then there is the spin difference. Defining the channel spin  $s$  as the resultant of the nucleus and neutron spins, we have the examples;

$${}^1\text{H} \quad a_{s=1} = 5.42 \text{ fm}$$

$$a_{s=0} = -23.75 \text{ fm}$$

$${}^2\text{H} \quad a_{s=3/2} = 6.34 \text{ fm}$$

$$a_{s=1/2} = 0.65 \text{ fm}$$

Thirdly, we have the possibility of sizeable imaginary components. The lightest nuclide in this class is  ${}^3\text{He}$ , with a bound atom coherent scattering length of

$$b_{\text{coh}} = 5.73 - i1.48 \text{ fm}$$

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All these features are consequences of nuclear excitation levels. Because the neutron energy range of interest to the condensed matter scientist is so tiny compared to the energy scale of the nuclear levels in these examples, there is no energy variation of any significance, but the isotope and spin variations are of major use.

The data base seems to be quite sound for many of the light nuclides. For heavier nuclides, however, there appear to be many gaps and anomalies, even before we begin to consider the consequences of the much more closely spaced nuclear levels in these systems. As an example of the cause for concern about the available data base, I show a table of the data for samarium and its isotopes as it could well appear in the kind of compilation easily available to the neutron scatterer at the present time (this Table is only slightly disguised from an actual one).

Table 1. *State of the Nuclear Data Base for heavy nuclides*

Typical case: Samarium and its isotopes

Isotope	$a$ (fm)	$\lambda$ (Å)	Comment
$^{152}\text{Sm}$	$0.7\pm 0.2$	1.86	From measurement <sup>1</sup>
	$0.0\pm 0.05$	1.30	From measurement <sup>2</sup>
	$-0.12\pm 0.04$	1.27	From measurement <sup>3</sup>
	$4.2\pm 0.3$		From compilation <sup>4</sup>
	1.4	1.80	From compilation <sup>4</sup>
	$2.9\pm 0.06$	1.30	Meas. <sup>2</sup> (Imag.comp.)
	-1.5	1.80	Comp. <sup>4</sup> (Imag.comp.)
$^{144}\text{Sm}$	$4.4\pm 4.0$		From compilation <sup>3</sup>
$^{147}\text{Sm}$	$14.0\pm 3.0$		From compilation <sup>4</sup>
$^{148}\text{Sm}$	$33.0\pm 6.0$		From compilation <sup>3</sup>
$^{149}\text{Sm}$	$-24.0-i11.0$	1.80	From compilation <sup>4</sup>
	$-18.7\pm 0.28$	1.80	From measurement <sup>6</sup>
$^{150}\text{Sm}$	$14.0\pm 3.0$		From compilation <sup>4</sup>
$^{152}\text{Sm}$	$-5.0\pm 0.6$		From measurement <sup>7</sup>
$^{156}\text{Sm}$	$9.25\pm 1.0$		From compilation <sup>4</sup>
	$8.0\pm 1.0$		From measurement <sup>7</sup>

We see from this Table that, firstly, the entries in commonly used compilations do not necessarily agree, even qualitatively, with measured and well-documented values. Secondly, there are a number of nuclides for which directly measured scattering lengths do not exist, yet values are given in compilations; in the compilations referred to above, no argument or evidence is advanced for the

numbers quoted. There appear to be almost 60 elements and isotopes for which only this kind of value is available.

Thirdly, there is a sign discrepancy for the imaginary components of the scattering length of the natural element. There has been a degree of arbitrariness in the literature over the choice of sign for the imaginary part; subsequently in this report, I employ the negative sign that comes naturally from the deduction of the scattering length from nuclear reaction theory (see also Sears' on this point).

Finally, there is a wave-length variation in the scattering for natural samarium, which is due to a strong resonance in  $^{149}\text{Sm}$  at low energy. In compilations, this wave-length variation is seldom made very obvious, even though in a few cases, such as natural samarium and its resonant isotope, a series of measurements from long wavelength to just below  $1\text{\AA}$  have been made<sup>3,6</sup>. It is my main purpose in this report to show how existing cross-section data can be used to give better scattering length data, including especially the systematic energy variation and phase component that are important for very many heavy nuclides.

## 2. Theory of neutron scattering lengths

The expansion in spherical polar coordinates of a plane wave travelling in the  $z$  direction with velocity  $v$  is, ignoring neutron and target spin effects,

$$\exp(ikz) = \frac{\sqrt{\pi}}{kr\sqrt{v}} \sum_{\ell=0}^{\infty} i^{\ell+1} [I_{\ell}(kr) - O_{\ell}(kr)] Y_{\ell 0}(\theta, \varphi) \quad (1)$$

Here,  $k$  is the neutron wave number,  $Y_{\ell 0}(\theta, \varphi)$  are spherical harmonics and  $I, O$  are combinations of spherical bessel and neumann functions and have the form of incoming and outgoing waves. Ignoring the centre-of-mass factor for the heavy nuclei considered in this report  $k$  has the numerical value

$$k = 0.0021968 \sqrt{E}$$

in units of  $10^{12} \text{ cm}^{-1}$ , where  $E$  is the neutron energy in electron-volts (eV).

For  $\ell = 0$ ,  $I_0 = \exp(-ikr)$ ,  $O_0 = \exp(ikr)$ , so we can write

$$\exp(ikz) = (\sqrt{\pi}/kr\sqrt{v}) [2\sin(kr)/\sqrt{4\pi} + \sum_{\ell \geq 1} i^{\ell+1} (I_{\ell} - O_{\ell}) Y_{\ell 0}(\theta, \varphi)] \quad (2)$$

For  $kr \ll 1$ ,  $\ell \geq 1$ ,  $I - O$  is negligible, so the effect of a nuclear potential confined to a region with radius  $R$  satisfying the first condition on these waves of higher angular momentum is unimportant for the scattering of the plane wave.

For a hard sphere of radius  $R$ , the wave-function has a node at  $R$  and is zero within the sphere. Hence the form of the wave-function of the impinging projectile plus scattering is

$$\Psi \propto \sin[k(r-R)]/kr \quad (3)$$

This can be written in the form of incoming and outgoing waves as

$$\Psi \propto \exp(ikR) [\exp(-ikr) - \exp(-2ikR)\exp(ikr)] \quad (4)$$

The amplitude of the scattered wave [ $\alpha (\sin k(r-R) - \sin kr)$ ] is approximately proportional to  $R$ , showing that this is the scattering length for a hard sphere. More accurately, the proportionality of the scattered wave amplitude can be written as  $[1 - \exp(-2ikR)]$ .

The physical situation described above is that in which the only channel is the elastic one. Here I define the concept of a channel. A channel is the distinct separation of a reacting nuclear system into two entities, one a residual (or target) nucleus, the other an ejectile (or projectile), each in a particular state of excitation and spin. Neutron cross-sections can be expressed in terms of a collision matrix  $U$ , which is the set of amplitudes of the outgoing wave-functions produced in the exit channels as a result of a neutron reaction with a target nucleus. For many purposes the most significant of these amplitudes is that for elastic scattering, from which can be deduced the elastic scattering cross-section, the total cross-section and the absorption cross-section (the sum of all cross-sections into non-elastic channels), and I shall limit our discussion henceforth to this one amplitude, which I shall denote by  $U$ .

In a plane wave the amplitude of the outgoing wave-function in the elastic channel relative to that of the incoming wave is unity. The form of the scattering wave-function is therefore proportional to  $(1-U)$ . In the example above  $U$  is just  $\exp(-2ikR)$ . For a potential well, rather than a rigid sphere, the quantity  $R$  is replaced by an apparent "radius"  $R'$ , which has the form illustrated in Fig.1; this is actually the phase shift (divided by  $k$ ) for the elastic scattering. This quantity can be visualised as the linear extrapolation of the sine wave outside the well to its node. This is illustrated in Fig.2 in which a resonance condition is achieved when the wave function inside the well contains exactly an odd number of quarter wavelengths and an infinite scattering length ensues. At an energy slightly below resonance the extrapolated value of the node (scattering length) is small, or even negative, while above resonance it is positive and much larger than the well radius. As a function of projectile energy the apparent radius for a specific potential well will have the form shown in Fig.3.

Translated into nuclear physics, in which the nuclear radius is approximately proportional to the cube root of the nuclear mass number  $A$ , the phase shift has the mass number dependence shown in Fig.4, the strong swings occurring as the potential well radius passes through the resonance condition for very low neutron energy, and it is found that nuclear scattering lengths do have a very rough and ready mass number dependence of this kind. Again translated to nuclei, the resonance effects apparent in Fig.3 would have spacing on a scale of several MeV.

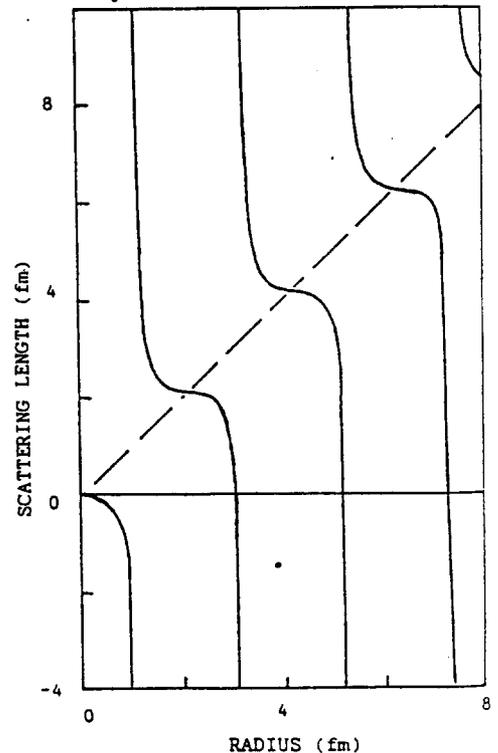


Figure 1. Scattering length as function of radius of a potential well. The broken line is for an impenetrable sphere.

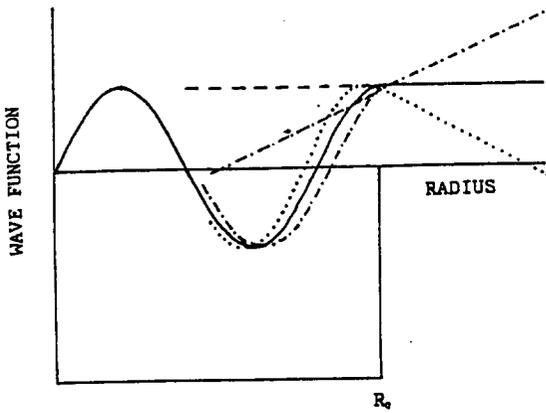


Figure 2. Wave function at resonance (full curve), below resonance (dot-dash) and above resonance (dotted). The node of the wave outside the well radius (or its extrapolation inside) is the scattering length.

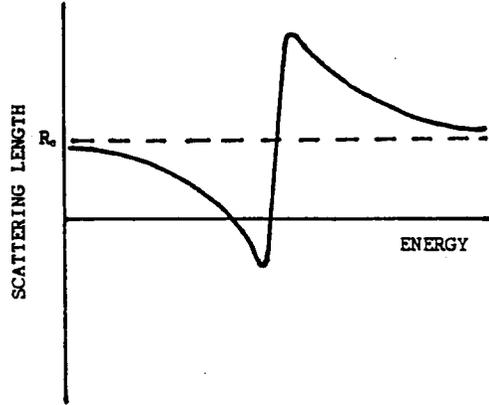


Figure 3. Scattering length as function of energy. At non-zero energy the swing is curtailed.

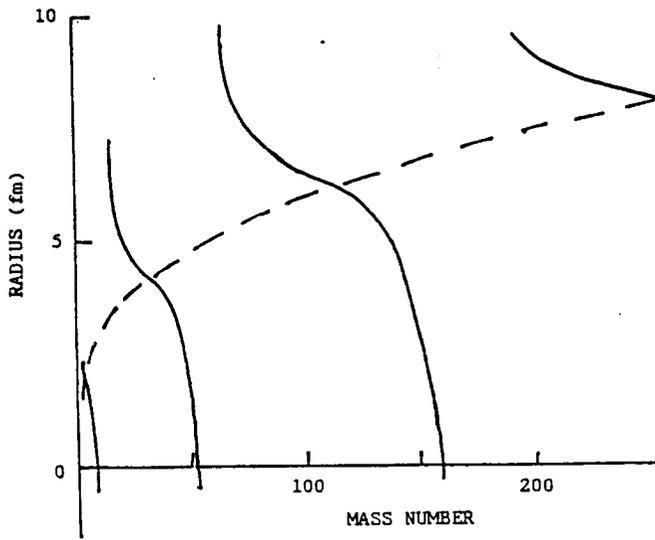


Figure 4. Scattering length as function of mass number.

Real nuclei are much more complex in their reactions with particles than can be represented by a simple potential well. The complexity is illustrated in Fig.5, which shows the energy levels of the compound nucleus formed by adding a neutron to a target nucleus. The binding energy of a neutron in the compound nucleus is several MeV. At this excitation energy, reached by adding a very slow neutron to the target, the level spacing is several tens of keV for very light nuclides, a few keV for medium mass nuclides, and a few eV or less for very heavy nuclides. It is these levels close to zero neutron energy, the resonances and the weakly bound states, that have a major influence on the scattering length. The reason why there is still an apparent potential well effect in the overall pattern of scattering lengths is that the compound nucleus levels retain some 'memory' of the potential resonances, which then exerts a long range influence on the zero energy properties of the scattering.

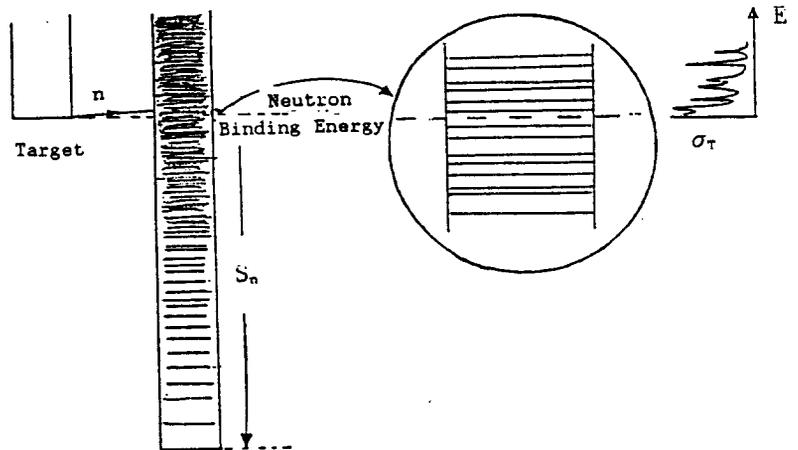


Figure 5. Nuclear energy levels of compound system. Inset shows enlargement near neutron binding energy with resulting cross-section.

For s-wave neutron interaction, the detailed form of the scattered wave is

$$\psi^{\text{scat}} = \frac{i/\pi}{kr/v} \sum_{s=|j-1/2|}^{j+1/2} [1 - U_s] \exp(ikr) \zeta_s Y_{00}(\vartheta, \varphi) \quad (5)$$

Here I have introduced the concept of the channel spin  $s$ , the total spin formed from the neutron spin  $\sigma (=1/2)$  and the spin of the target nucleus  $j$ ; the normalised channel spin wave-function thus formed is denoted by  $\zeta_s$ . When this is compared with the usual form for the elastic scattering wave-function

$$\psi^{\text{scat}} = \sum_{s=|j-1/2|}^{j+1/2} f_s \exp(ikr) \zeta_s / r \quad (6)$$

the relation

$$f_s = i(1 - U_s)/2k \quad (7)$$

is obtained.

The usefulness of this expression lies in the fact that a full formal understanding of the collision matrix element can be obtained from the R-matrix nuclear reaction theory of Wigner and Eisenbud<sup>8</sup> (see also Lane and Thomas<sup>9</sup>). For thermal neutron reactions with nearly all nuclides except the actinides, elastic scattering is accompanied by radiative capture and alpha-emission channels with partial widths that are very small compared with the resonance spacing. For such

reactions an excellent approximation that can be derived from the full theory is the reduced R-matrix formulation. In this, the collision function for elastic scattering can be expressed as

$$U = e^{-2ika}[1 - ika\mathfrak{R}]^{-1}[1 + ika\mathfrak{R}] \quad (8)$$

where the quantity  $a$  is a channel radius (which is usually chosen to equal either the potential well radius of the nucleus or the potential scattering length  $R'$ ). The reduced R-function for this one-channel case is

$$\mathfrak{R} = \sum_{\lambda} \frac{\gamma_{\lambda(n)}^2}{E_{\lambda} - E - i\Gamma_{\lambda(a)}/2} \quad (9)$$

Here, the  $E_{\lambda}$  are the formal nuclear energy levels of the (compound) system, with reduced neutron width amplitude  $\gamma_{\lambda(n)}$  and absorption width  $\Gamma_{\lambda(a)}$ , which is the sum of partial widths of all reaction widths of the compound nucleus level  $\lambda$ . The functions  $U$  and  $\mathfrak{R}$  are implicitly labelled by the channel spin  $s$ , as are the nuclear levels  $\lambda$ .

If the nucleus can be represented by a simple potential well, the levels  $\lambda$  are just those causing the fluctuations  $R'$  apparent in Fig.3 (or 2). If there are no compound nucleus levels (resonances) with  $E_{\lambda}$  close to  $E$ , the effect of these potential well levels can be incorporated into the effective channel radius  $a$ , which then equals the potential scattering length  $R'$ . The quantity  $ka\mathfrak{R}$  is then negligible, and since  $ka \ll 1$  for low neutron energy, the scattering amplitude  $f$  is approximately equal to  $-a$ , which is the common usage for most neutron scattering work. If compound nucleus levels are significant but distant, so that  $\mathfrak{R}$  is effectively real and hardly changes over the energy range of interest, the relation is  $f \approx -a(1-\mathfrak{R})$ .

These relations are for a single channel spin. In general  $a$  and  $\mathfrak{R}$  will vary for different isotopes and channel spins. For the general case of a multi-isotope element with non-zero-spin isotopes, the coherent scattering length (for free nuclei) has the well-known form

$$a_{\text{coh,el}} = -\sum_i w_i \sum_s g_{i,s} f_{i,s} \quad (10)$$

where  $w_i$  is the fractional abundance of the isotope  $i$  in the element and  $g_{i,s}$  is the spin-weight factor  $(2s + 1)/[2(2j_i + 1)]$  for channel spin  $s$ . The total scattering cross-section is

$$\sigma_{sc} = (\pi/k^2) \sum_{i,s} w_i g_{i,s} |1 - U_{i,s}|^2 \quad (11)$$

the absorption cross-section is

$$\sigma_a = (\pi/k^2) \sum_{i,s} w_i g_{i,s} (1 - |U_{i,s}|^2) \quad (12)$$

and the total cross-section is

$$\sigma_T = 2(\pi/k^2) \sum_{i,s} w_i g_{i,s} (1 - \text{Re}U_{i,s}) \quad (13)$$

For many of the heavy nuclides, a single resonance level dominates the energy variation of the collision function for given channel spin. I give here

the single level approximation for the scattering amplitude, with an important extension to include the milder energy dependent effects from more distant resonance terms. This is obtained from eqs.7,8,9. The expression for the scattering amplitude  $f_{i,s}$ , to second order in  $a$ ,

$$f = -[a(1-R_{ex})] + 2ka^2S_{ex}(1-R_{ex}) + \frac{a\gamma_{\lambda(n)}^2[(E_{\lambda}+\Delta_{\lambda}-E) + ka\Gamma_{\lambda}(1-R_{ex})]}{(E_{\lambda}+\Delta_{\lambda}-E)^2 + \Gamma_{\lambda}^2/4} \\ + i \left\{ [aS_{ex} + ka^2(1-R_{ex})^2 - ka^2S_{ex}^2] + \frac{a\gamma_{\lambda(n)}^2[\Gamma_{\lambda}/2 + 2ka(E_{\lambda}+\Delta_{\lambda}-E)(1-R_{ex})]}{(E_{\lambda}+\Delta_{\lambda}-E)^2 + \Gamma_{\lambda}^2/4} \right\} \quad (14)$$

the labels  $i,s$  being implicit. Here, the quantities  $R_{ex}$ ,  $S_{ex}$  are measures of the real and imaginary contributions of other levels; when these are some distance from the level  $\lambda$  the distant level quantities can be described as simple polynomials:

$$R_{ex} = A_r + B_r(E - E_0) + C_r(E - E_0)^2 + \dots \quad (15)$$

$$S_{ex} = A_i + B_i(E - E_0) + C_i(E - E_0)^2 + \dots \quad (16)$$

where  $E_0$  is some approximately central point in the energy range of interest. The width  $\Gamma_{\lambda}$  is the total resonance width of the most significant resonance  $\lambda$ ; it is the sum of a neutron width  $\Gamma_{\lambda(n)}$  ( $= 2ka\gamma_{\lambda(n)}^2$ , to second order in  $ka$  and distant level quantities) and the absorption width  $\Gamma_{\lambda(a)}$ . The quantity  $\Delta_{\lambda}$  is a level shift that can usually be neglected at the very low neutron energies considered in this paper.

The generalised single level formula of eq.14 can be used to analyze cross-section and scattering length data by least square fitting to obtain an assessment of the coherent scattering length data from neutron energy close to zero up to approximately 0.5 eV. An alternative way of treating the potential well effects of Figs.1-3 is to incorporate their long-range effects empirically in the constant  $A_r$  of eq.11, and to use the conventional nuclear potential radius as the channel radius  $a$ :

$$a = 1.16A^{1/3} + 0.6 \text{ fm.} \quad (17)$$

The pivot energy  $E_0$  is also chosen in advance. With these constraints the parameters  $E_{\lambda}$ ,  $\gamma_{\lambda(n)}^2$ ,  $\Gamma_{\lambda(a)}$ ,  $A_r$ ,  $B_r, \dots$ ,  $A_i$ ,  $B_i, \dots$  for each isotope and channel spin are then determined from the data. The free atom nuclear coherent scattering lengths can then be determined from eqs.10 and 6. Because there is generally very much more detailed cross-section data available on a given nuclide than data on scattering lengths, it is possible to learn a great deal about the latter by this procedure.

### 3. Examples of deduced scattering lengths

The main example that I shall use to demonstrate the possibility of extraction of the energy dependent, complex scattering length from cross-section data is the case of samarium and its isotopes, also discussed in the introduction and Table 1.

The total cross-section of natural samarium<sup>10</sup> is shown in Fig.6. The strong resonance at approximately 0.1 eV is known (from measurements on separated isotopes) to belong to the target <sup>149</sup>Sm (present at 13.9% in the natural element), and the spin  $s$  of this resonance has been ascertained by polarised neutron and target techniques<sup>11</sup> to be  $s = 4$  ( the target spin  $j = 7/2$ ). The resonance parameters have been rather well determined<sup>12</sup>. The scattering cross-section as a function of energy across the resonance has been measured<sup>13</sup>. The refitting of all these data gives a set of resonance parameters rather close to those found in the literature together with parameters for the slowly varying contribution of the other levels. From these, the coherent scattering length behaviour of <sup>149</sup>Sm is calculated; it is shown as the broken curve in Fig.7.

The data points shown in Fig.7 have not been used in the fitting. They are measurements of the real part of the neutron scattering length of <sup>149</sup>Sm, made by interferometry techniques<sup>6</sup>, and are not generally found in compilations. It is clear that the broken curve represents the data rather well but not optimally so. A refit of the collision function parameters to include these data leads to the calculated full curve, which is obviously better, but not drastically so. It is clear that it is possible to calculate the energy dependence of coherent scattering lengths quite reliably from cross-section data of this quality.

If we now use the deduced parameters of the collision function of <sup>149</sup>Sm with the measured long wave-length coherent scattering lengths of <sup>152,154</sup>Sm and the natural element we can determine the energy dependent scattering length of natural Sm and the abundance-weighted average of all the

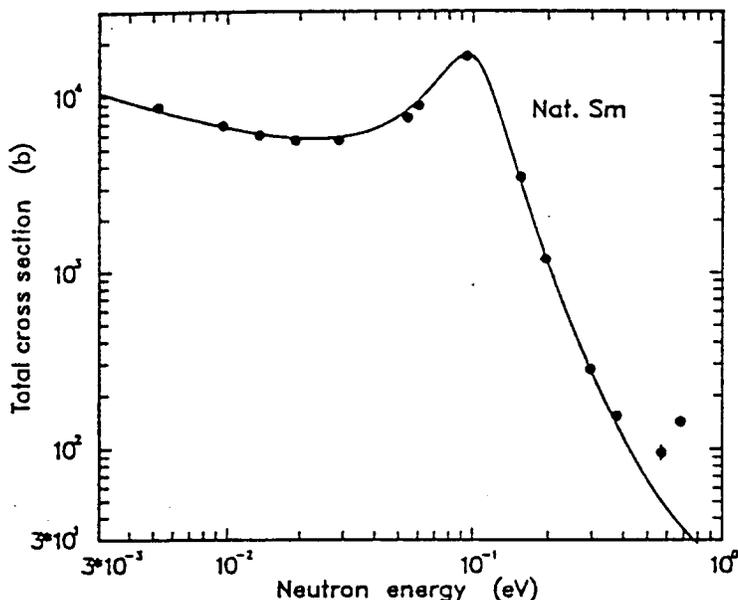


Figure 6. The neutron total cross-section of natural samarium.

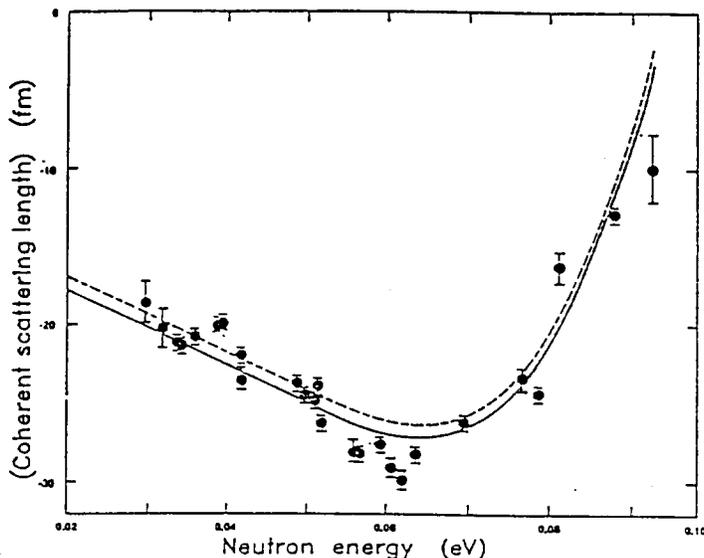


Figure 7. Real part of the coherent scattering length of <sup>149</sup>Sm. Curves and data are explained in the text.

isotopes except 149, 152, and 154. The coherent scattering length of the natural element is shown as the full curve of Fig.8. This can be compared with some energy dependent measurements<sup>7</sup> which have not been used in the fitting procedure. It is clear that the agreement is very good, again giving us confidence in our procedure.

The analysis does not give us confidence in the overall set of numbers in Table 1, however. The weighted average of the scattering lengths of the isotopes for which measurements do not seem to be available is determined in our fitting as 8.5 fm. The value computed from the entries of Table 1 is 19 fm. The discrepancy is wide, and a large part of it comes from the extraordinarily high value given in ref.5 for <sup>148</sup>Sm. If this were really so large, it would have to be due to a very strong bound level in the <sup>148</sup>Sm + n system, and no evidence in the form of a very high total scattering or capture cross-section is available to support this hypothesis.

Phil Seeger and I have carried out analysis of this kind on the available cross-section data of all the elements of the rare earth series that have obvious resonance effects<sup>14</sup>. In general we have taken our analysis and computation of coherent scattering lengths up to 0.5 eV (wavelength 0.41 Å). The calculated, and recommended, energy dependence for all the isotopes and some natural elements that display resonance behaviour is shown in Fig.9. We plan to make a similar systematic study of other nuclides at a later date.

#### 4. Some effects of resonance behaviour in scattering lengths.

##### 4.1. Small angle scattering and total reflection

The striking features of the coherent scattering lengths in the examples shown in Fig.9 are of course the large imaginary component and the very large real component above the resonance energy. The former gives rise to strong absorption; this severely limits the penetration depth of neutrons, but also gives a phase variation that can yield additional information from scattering measurements.

<sup>115</sup>In has a strong resonance at 1.46 eV. From the resonance parameters the scattering length can be calculated. Its energy behaviour above 1 eV is shown in Fig.10. Experiments have been carried out to detect the anomalous total scattering caused by this<sup>15</sup>. The calculated scattering from a plane surface of indium at various neutron energies is shown in Fig.11. There is a striking variation in the angle of quasi-critical-scattering with a maximum reached at about 1.6 eV, and this is greater than at any energy above 0.1 eV

Resonance critical scattering is probably observable at energies well above 1 eV for several materials. An example is shown for <sup>152</sup>Sm, which has a particularly strong resonance at 8 eV, in Fig.12.

An interesting example of the possible application of resonance scattering is being explored by Hjelm, Seeger and Thiagaryan<sup>16</sup>. They have suggested the use of the strong contrast provided by the large change in coherent scattering length over a small range of neutron energy (e.g. 63 fm from 0.07 eV to 0.11 eV in <sup>149</sup>Sm) to locate the position of the metallic atom in large metallo-organic molecules, using small angle scattering diffractometers on white neutron pulsed sources.

## 4.2. Bragg reflection

Despite the strong absorption, strong Bragg reflections can be obtained at neutron energies close to resonance. One of the strongest resonances known at low energies is found in the cross-section of  $^{240}\text{Pu}$  at 1.04 eV. A calculated Bragg reflection from a single crystal of  $\text{PuO}_2$  is shown in Fig.13. This is the [12,4,0] reflection at the energy 1.04 eV. It is calculated for a crystal thickness of  $2.10^4$  atomic layers, by using the Darwin multiple scattering formalism originally developed for X-ray diffraction. The angle-integrated line strength over the energy range 0.9 eV to 1.2 eV is shown in Fig.14. For this thickness of crystal the reflection strength qualitatively follows the resonance scattering cross-section. For increasing crystal thickness the wings rise, the centre remaining almost constant owing to primary extinction and absorption, until for very great thickness there is little evidence of the resonance.

These general features suggest that resonance scattering could be used to explore the structure of surface layers of intermediate thickness, although not as low as a few mono-layers. It should also be remembered that, for nuclides with odd numbers of protons, neutrons or both the resonance scattering is sensitive to the relative spin polarisation of the neutron beam and the target nucleus, and could therefore be useful in studies of magnetic materials at low temperatures.

## 4.3. Powder diffraction

Because of the strong absorption resonance scattering would not appear at first sight to be a useful effect in powder diffraction studies. Nevertheless, in refinement analyses of the Reitveld kind it is obviously important to include as well as possible the wave-length variation of the scattering length, and the changing ratio of the imaginary to real components offers the opportunity to gain some phase information.

Even relatively weak resonances complicate the analysis of a powder pattern. An example, due to Goldstone and Lawson<sup>17</sup>, is shown in Fig.15; this is  $\text{NpD}_{2.67}$ , which, at room temperature has a face-centred cubic cell, with the deuterium atoms occupying tetrahedral and octohedral sites in different proportions. The absorption dips in the overall scattering (coherent and incoherent) due to the resonances at 0.49 eV and 1.32 eV are very strong. Greater detail for the resonance region is shown in Fig.16. In this some evidence of diffraction lines is to be found in the resonance minimum. Indeed, these, as well as the lines above the resonance (in energy), may be surprisingly strong relative to the lines at longer wavelengths below the resonance.

A preliminary analysis indicates that the relative magnitude of these groups of lines certainly cannot be explained on the assumption of a constant

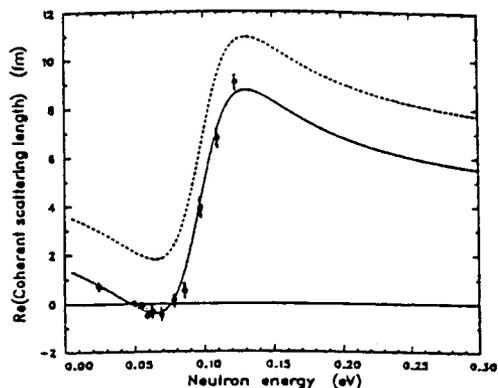


Figure 8. Real part of the coherent scattering length of natural samarium. Curves and data are explained in the text.

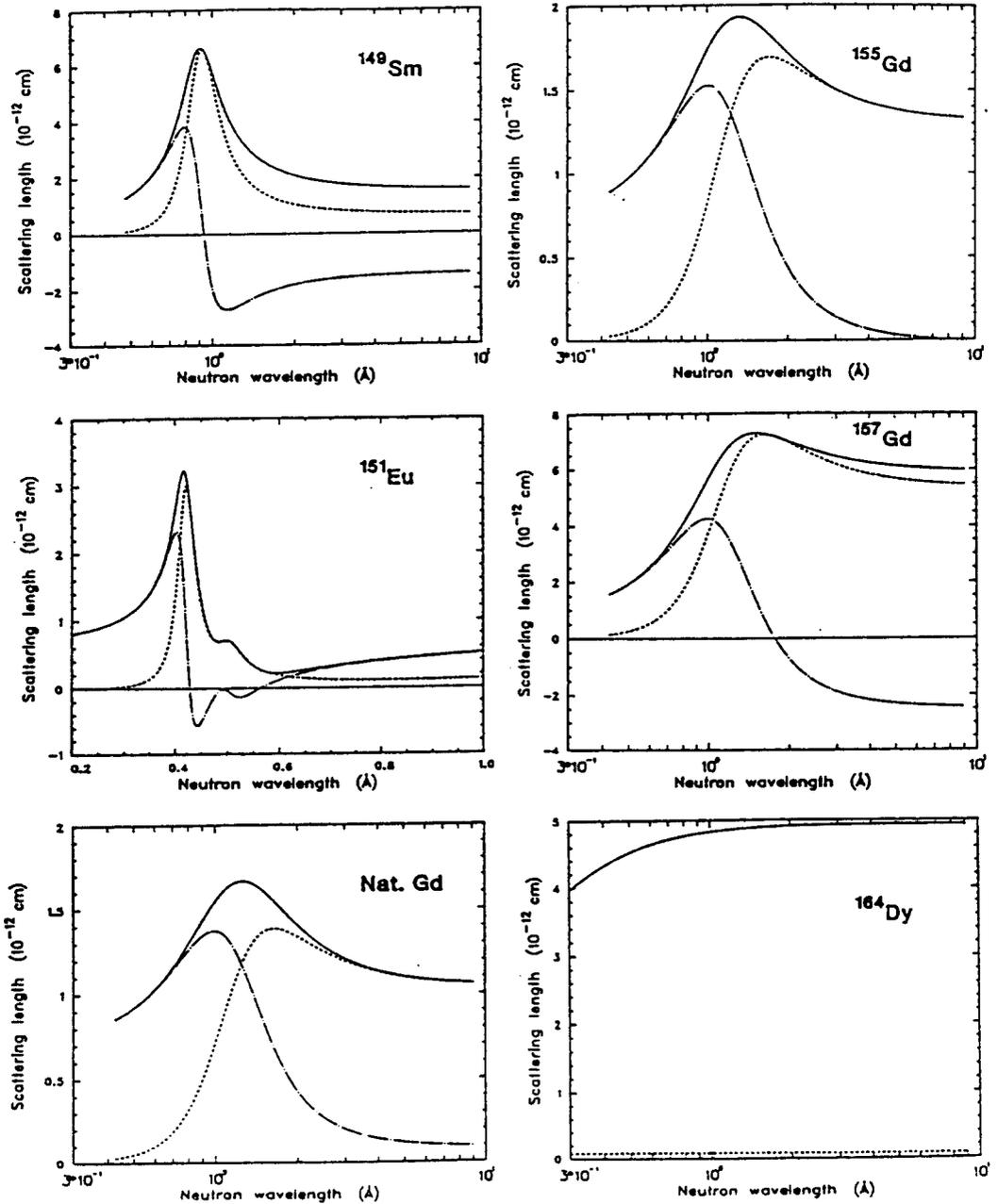


Figure 9. The coherent scattering lengths of the rare earth nuclides with resonance effects in the thermal and epithermal neutron energy region. The dash-dot curve is real, dot is imaginary, and full is the modulus.

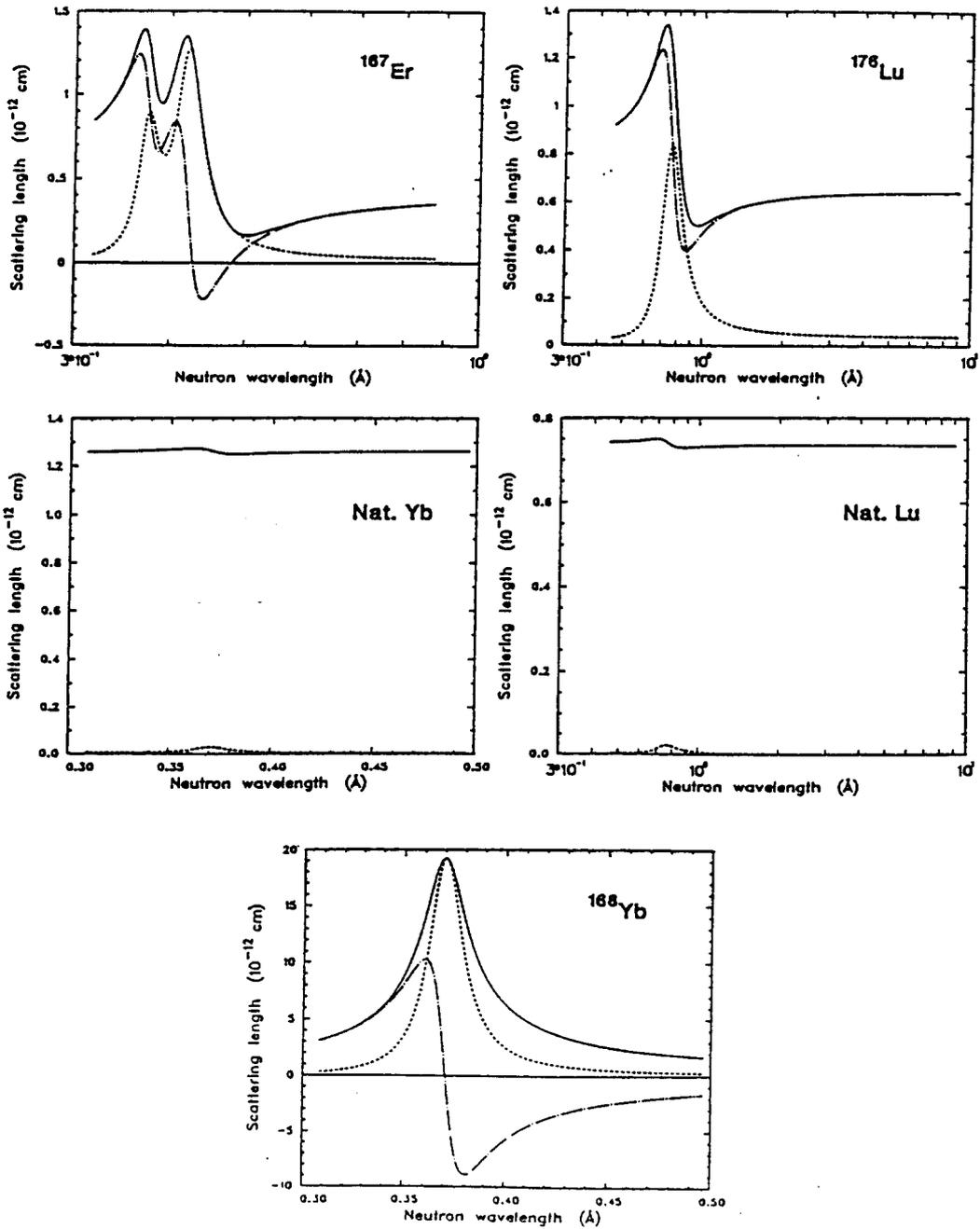


Figure 9. Continued

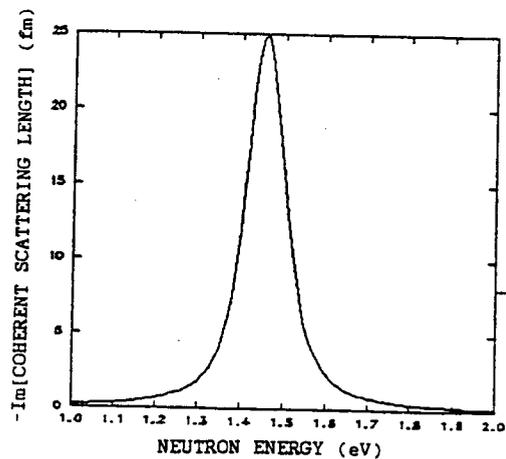
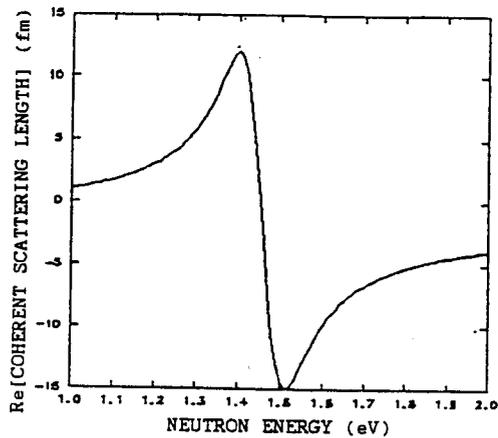


Figure 10. Real and imaginary scattering length of indium as function of neutron energy.

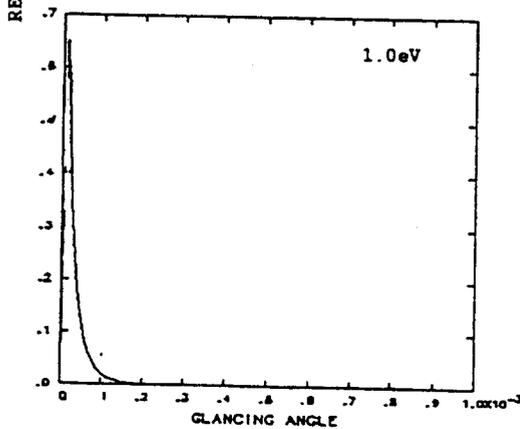
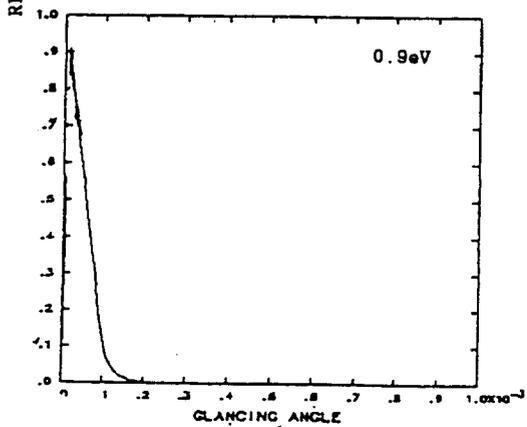
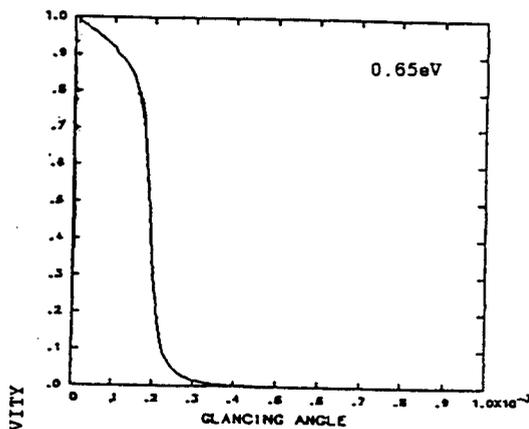
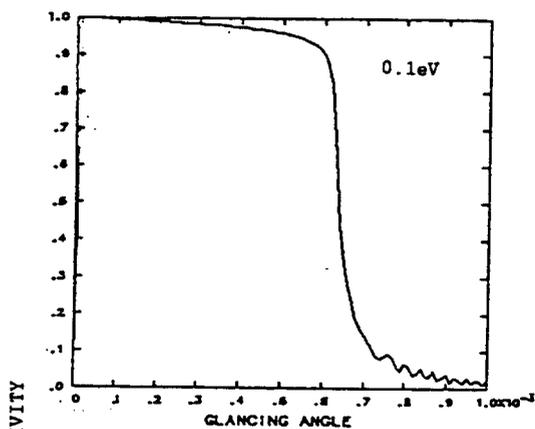


Figure 11. Reflectivity of indium as function of glancing angle for a range of neutron energies.

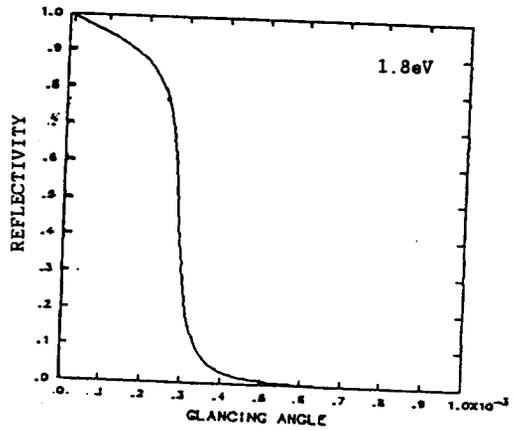
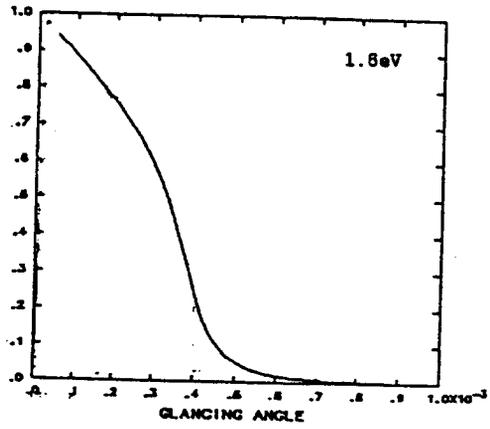
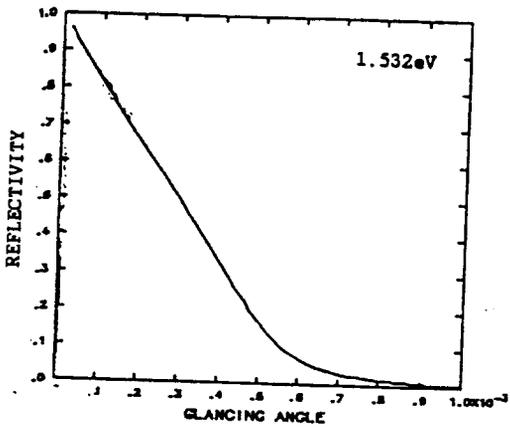
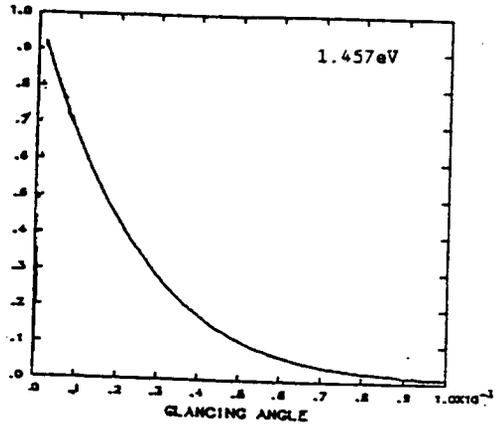
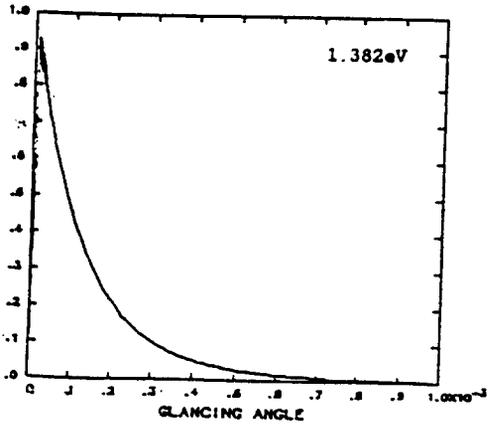


Figure 11. Continued

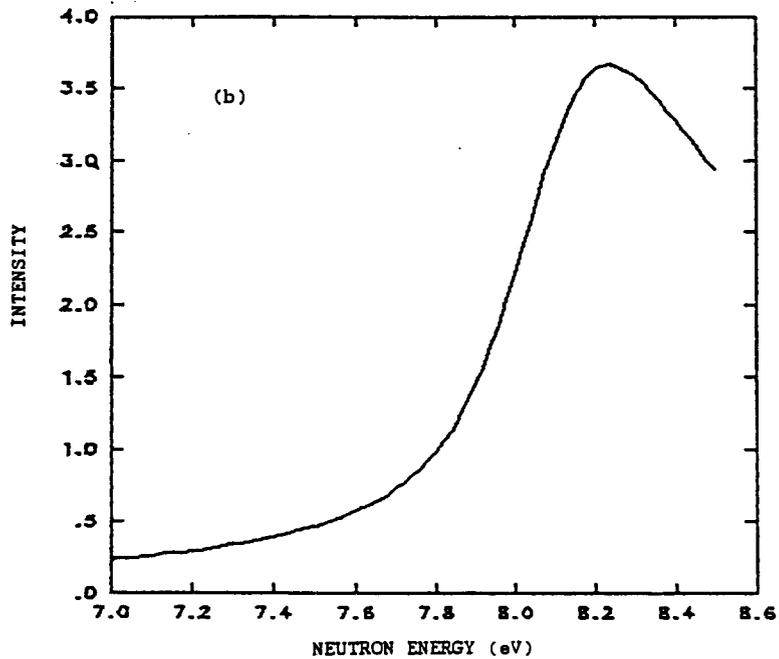
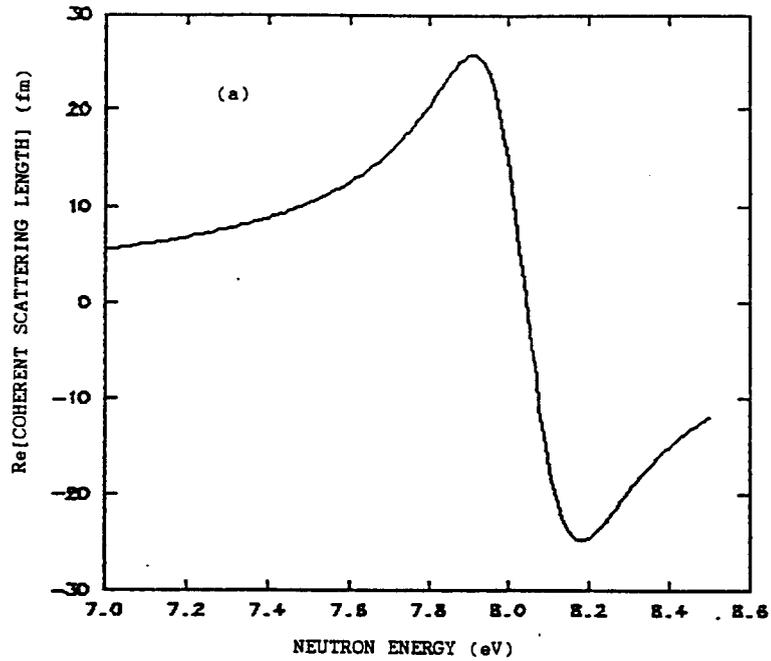


Figure 12. Critical reflection properties near the 8 eV resonance of  $^{152}\text{Sm}$ . Part (a) shows the real part of the scattering length. Part (b) shows the neutron reflection strength integrated from 0 to 0.1 mrad as function of neutron energy.

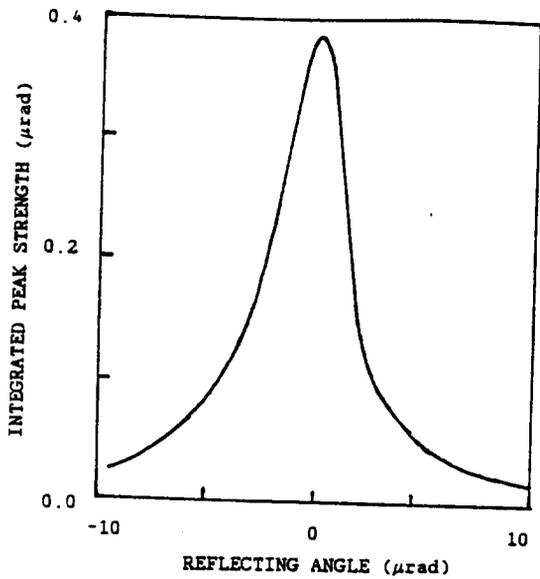


Figure 13. Detailed angular dependence of Bragg reflection line for  $^{239}\text{PuO}$ , for neutrons of 1.068 eV.

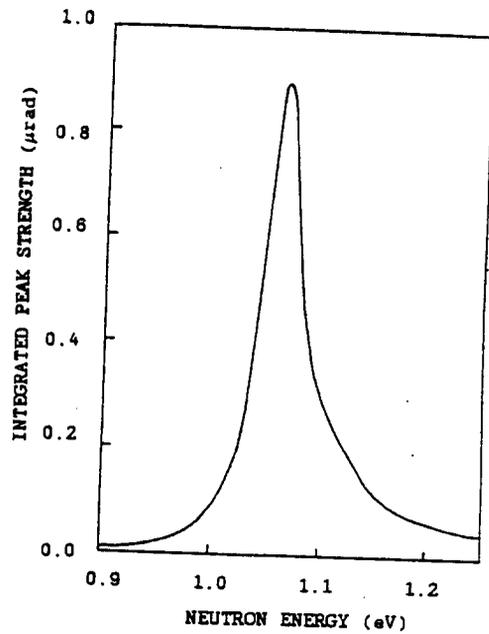


Figure 14. Energy dependence of Bragg reflection strength for reflection from  $10^{\circ}$  atomic layers as function of neutron energy near the 1.04 eV resonance of  $^{239}\text{Pu}$ .

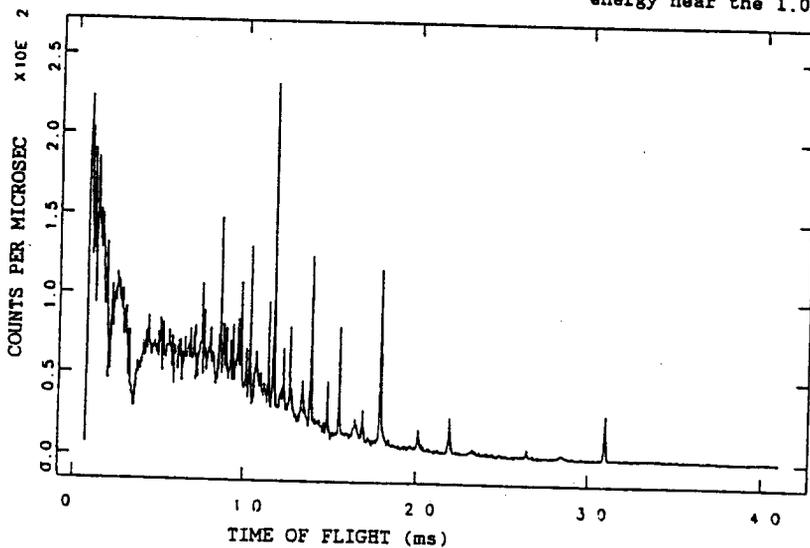
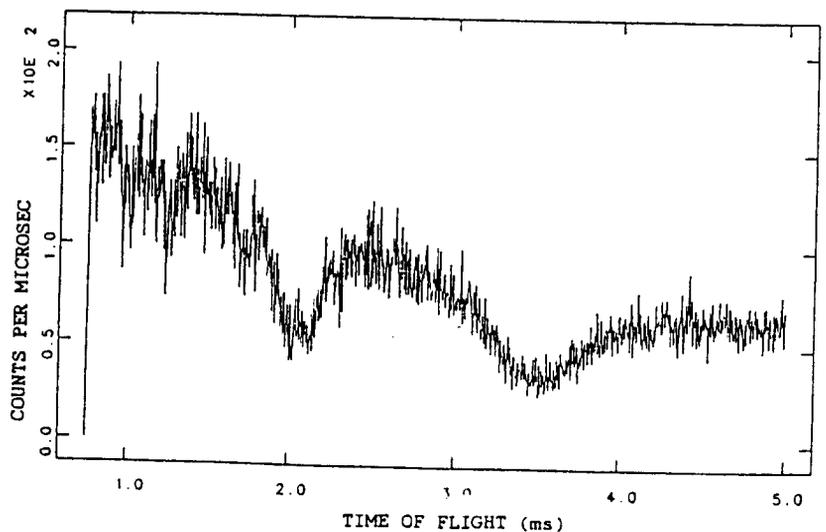


Figure 15. Powder diffraction pattern from  $\text{NpD}_{2.8}$ , as function of neutron time-of-flight.

Figure 16. Powder diffraction pattern from  $\text{NpD}_{2.8}$ , as function of neutron time-of-flight, showing resonance region in greater detail



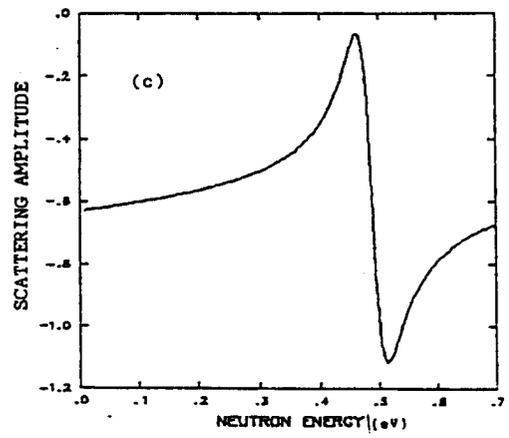
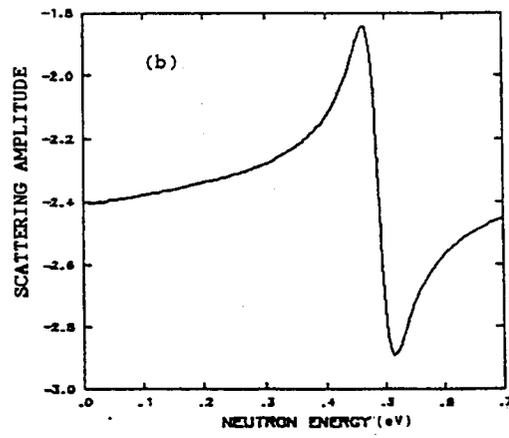
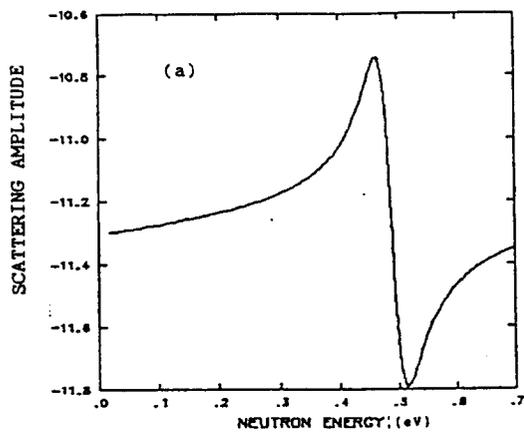


Figure 18. Diffraction line strengths for  $\text{NpD}_{1.8}$ , calculated as function of neutron energy in the neutron energy. Line classes as in Fig.17.

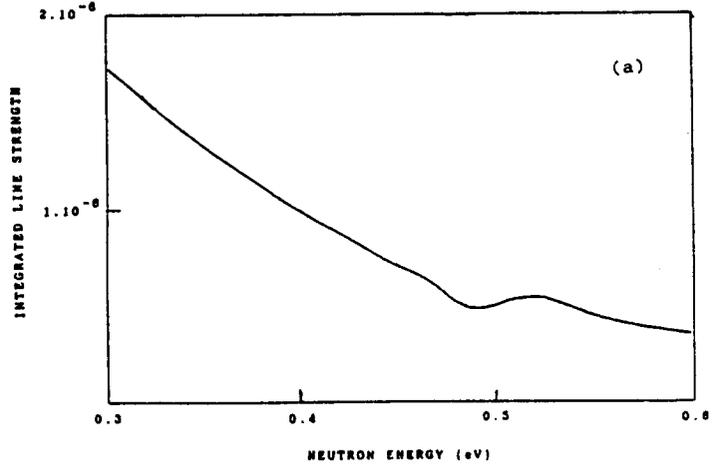
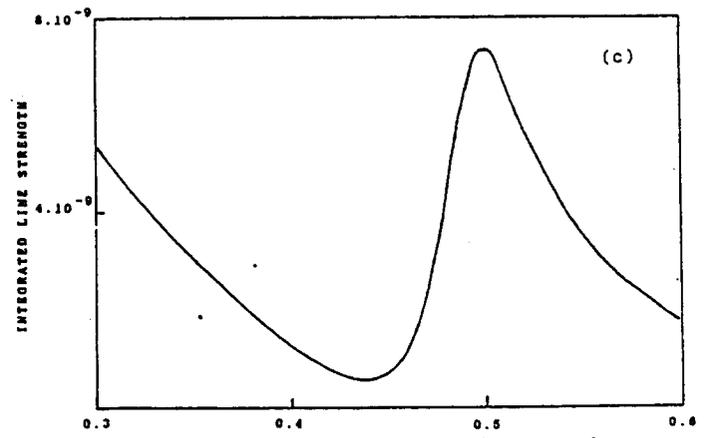
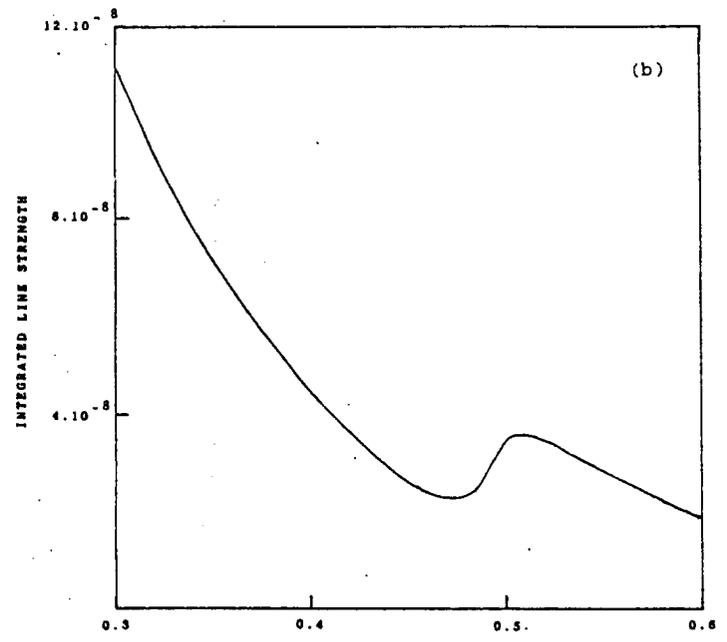


Figure 18. Diffraction line strengths for  $\text{NpD}_{1.8}$ , calculated as function of neutron energy in the neutron energy. Line classes as in Fig.17.

coherent scattering length for neptunium, or even with the inclusion of the resonance variation if the phonon frequency for the deuterium is assumed to be the same as that for Np. The calculated scattering amplitude for the latter assumption is shown for the three different classes of the order of reflection in Fig.17. The largest amplitude, in general, is for reflection order satisfying the relation  $h + k + l = 4n$  where  $n$  is an integer, and the next largest is that for  $h + k + l = 2n + 1$ . The corresponding energy dependence of the angle-integrated line strength is shown in Fig.18. It is clear that the contribution of deuterium is too high to allow the resonance effect for the  $h + k + l = 4n$  reflections to be very significant. Inelastic scattering measurements<sup>18</sup> have revealed, however, that very high frequency phonons reside on the deuterium, which will give it a small Debye-Waller factor, especially at the high energies corresponding to the resonance region.

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